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# **RESEARCH ARTICLE**

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#### **Key Points:**

- New and compiled paleomagnetic data from Mesozoic rocks of the Wrangellia superterrane show consistent inclinations and declinations
- Two new tectonic scenarios describe Wrangellia's drift from a Triassic subequatorial to a mid-Cretaceous, mid-latitude accreted position
- Robust and reproducible paleomagnetic data of the Wrangellia superterrane are useful for constraining Mesozoic plate motions of eastern Panthalassa

#### **Supporting Information:**

Supporting Information may be found in the online version of this article.

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# Paleolatitudinal Drift and Major Rotation of the Wrangellia Superterrane in the Mesozoic: A Signal of East-Panthalassa Plate Motion?

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Abstract The allochthonous origin of the Wrangellia superterrane relative to North America (NAM) has been established in the early days of plate tectonics using paleomagnetic and geologic data. However, orogenic deformation of the Wrangellia superterrane after its Cretaceous accretion to NAM has complicated the reconstruction of its pre-accretionary tectonic evolution. In particular, vertical-axis rotations linked to strike-slip faulting have cast doubt on the usefulness of paleomagnetic declinations in deciphering its pre-accretionary motions. Here, we compile paleomagnetic data from the Wrangellia superterrane and present new results from uppermost Triassic limestones and lowermost Jurassic lavas of the Bonanza arc, which confirm that it was at those times at a much lower latitude than today ( $\sim$ 50°-60°N), at a latitude of either  $\sim$ 25°-30°N or  $\sim$ 25°-30°S. Moreover, declinations reveal a coherent, major clockwise or counterclockwise rotation, depending on hemispheric origin. When correcting for previously documented true polar wander at the approximate longitude of the Wrangellia superterrane at ~190 Ma, new and existing paleomagnetic data allow for two possible scenarios of Mesozoic kinematic evolution: from 190 to 80 Ma, in the southern hemisphere scenario, the Wrangellia superterrane was transported  $\sim$ 5,000 km northward while rotating  $\sim$ 110° clockwise with an associated north-dipping subduction zone, while in the northern hemisphere scenario it remained at northern middle latitudes while rotating  $\sim 70^{\circ}$  counterclockwise at a south-dipping subduction zone. The southern hemisphere scenario appears more plausible when tested against recent reconstructions of minimum oblique convergence beneath NAM and mantle tomography images of slab remnants.

#### 1. Introduction

The Cordilleran orogen of northwestern North America (NAM) is among the largest accretionary orogens that formed during and since Mesozoic times. The orogen developed at a series of subduction zones that collectively accommodated plate convergence between the major Farallon and several smaller, mostly oceanic Panthalassa plates, and the largely continental NAM Plate (Dickinson, 2004; Engebretson et al., 1985; Nokleberg et al., 2000; Sigloch & Mihalynuk, 2013). Kinematic reconstruction of the Cordilleran orogen of NAM is important for connecting surface geologic processes to mantle dynamics (Sigloch & Mihalynuk, 2017; van der Meer et al., 2010, 2012), exploring economic resources (Nokleberg et al., 2005), and modeling Earth's paleogeography (Scotese, 2021) and paleoclimate (Caruthers et al., 2022; Dal Corso et al., 2020). Restoring the kinematic history of the Cordilleran orogen requires quantifying the amount, timing, and direction of past displacement of the fault-bounded crustal fragments that compose the orogen. This appears at first glance straightforward: the Cordilleran orogen exposes terranes that sutured against the North American margin in the Mesozoic, and that were transported northward since Late Cretaceous times along margin-parallel strike-slip faults, accommodated by complex deformation and oroclinal bending in Alaska (e.g., Cowan et al., 1997; Hildebrand, 2013; Irving et al., 1996; Johnston, 2001). But reading the plate kinematic history that led to the modern orogen has proved challenging.

Orogenic deformation of these terranes has complicated the reconstruction of their pre-accretionary motions. In particular, doubt has been cast on the pertinence of using paleomagnetic data of segments of the North American Cordillera. These data are the most direct quantitative sources to reconstruct past plate motion history, in particular of the rock units that unequivocally underwent a plate motion history within the eastern Panthalassa Ocean, west of NAM, such as the "Wrangellia Superterrane" that is found from British Columbia to Alaska (Figure 1). Well-known problems include: (a) remagnetization of sedimentary and igneous rocks (e.g., Butler,



**Figure 1.** Tectonic setting of the Wrangellia superterrane and other geological records discussed in the text. (a) Tectonic map of the Cordillera of western North America (NAM) (modified after Colpron & Nelson, 2009). The paleomagnetic localities are those of the studies compiled in Figure 7 and in the Data Set S3 (Supporting Information S1). OPS, Ocean plate stratigraphy. (b) UUP07 P-wave tomographic model at 1,670 km depth (Amaru, 2007) showing the location of the Socorro slab (offshore western Mexico) and the outline of the Cordillera of NAM. Positive seismic wave-speed anomalies, such as the Socorro slab, are in blueish colors (up to +0.63%), whereas negative seismic wave-speed anomalies are in reddish colors (down to -0.63%).

Gehrels, & Kodama, 2001); (b) inclination flattening of sedimentary rocks causing underestimation of paleolatitude (e.g., Krijgsman & Tauxe, 2006); (c) the use of paleomagnetic data obtained from intrusive rocks (e.g., Beck & Noson, 1972; Rusmore et al., 2013), from which interpretation is complicated by a lack of control on paleohorizontal at the time of magnetization; (d) local vertical-axis rotations of blocks during strike-slip faulting making the evaluation of regional (plate) tectonic rotation difficult (e.g., Beck, 1976; Irving et al., 1996). Moreover, recent advances in statistical procedures of paleomagnetic data compilation (e.g., Gallo et al., 2023; Vaes et al., 2023) make it timely to test whether there are flaws in the reference apparent polar wander path (APWP) that quantifies the paleolatitudinal position of NAM (e.g., Kent & Irving, 2010; Torsvik et al., 2012).

In this paper, we re-assess the coherence and consistency of paleomagnetic data from the Wrangellia superterrane. We compiled paleomagnetic data from the Wrangellia superterrane, covering the Paleozoic to Cretaceous, and also compiled data from its post-accretionary Late Cretaceous and Cenozoic rock record. In addition, we collected a new, large paleomagnetic data set from Upper Triassic sedimentary rocks and Lower Jurassic Bonanza arc lavas of the southern Wrangellia superterrane. We use these data to test the reproducibility of existing paleomagnetic data from the youngest Mesozoic age interval for which previous data revealed a subequatorial position. We compare these data to an updated global APWP of Vaes et al. (2023) and evaluate how North American paleomagnetic data sets compare to the global compilation by using data comparison methods (Vaes et al., 2022) that overcome statistical issues of classic paleomagnetic approaches (Rowley, 2019).

In detail, we test whether the paleomagnetic inclinations, estimates of paleolatitudinal position through time, provide a coherent pattern. We also evaluate whether the paleomagnetic declinations, estimating vertical-axis rotations, provide a coherent motion history for the terrane as a whole. We test rates of reconstructed

northward motion of the Wrangellia superterrane against recent reconstructions of minimum oblique subduction components obtained from kinematic evidence from the Californian forearc ophiolites (Arkula et al., 2023) and from Ocean plate stratigraphy (OPS) (rock sequences scraped off subducted ocean floor, including pillow lavas, overlying (hemi)pelagic sedimentary rocks, and foreland basin clastics; Isozaki et al., 1990) accreted in the Californian Franciscan subduction complex (Alvarez et al., 1980; Courtillot et al., 1985; Tarduno et al., 1985, 1986). Finally, we evaluate whether reconstructed positions of the Jurassic Bonanza arc at the time of the end of subduction coincide with the presence of slab remnants in mantle tomography. The updated statistical paleomagnetic procedures and reference frame, as well as the multiple independent tests, will determine whether the extensive rock records of the Wrangellia superterrane dating back to the Paleozoic provide accurate information on the paleolatitudinal and rotational components of its pre-collisional plate motion history relative to NAM.

## 2. Geological Context

#### 2.1. The Wrangellia Superterrane

Extending from western Alaska to southeastern British Columbia, the Wrangellia superterrane is subdivided into the Peninsular, Wrangellia, and Alexander terranes (Figure 1a), which overall share a multi-stage history of arc magmatism that spans the Late Paleozoic and most of the Mesozoic (~360-100 Ma; Alberts et al., 2021; Nokleberg et al., 1994; Plafker & Berg, 1994). The amalgamation of the three terranes is thought to have occurred during the Paleozoic, prior to their Mesozoic accretion to NAM. East, north, and south of the Wrangellia superterrane are the Intermontane superterrane of the Canadian Cordillera and the Franciscan accretionary prism of Oregon and California, respectively, which both formed as part of the North American margin and consist of metamorphosed continental margin rocks, ophiolites, and accreted complexes that were amalgamated in a subduction zone between the Wrangellia superterrane and the North American continent since the Early Jurassic (Nokleberg et al., 2000; Wakabayashi, 2015). The Wrangellia superterrane is inferred to have been located in an upper plate position relative to a subduction zone that formed the Bonanza arc, but in a downgoing plate position relative to NAM, until its collision with NAM, sometime in the middle to Late Cretaceous (Nokleberg et al., 1994, 2000; Plafker et al., 1989; Tikoff et al., 2023). Such a Mesozoic double subduction system is consistent with slab remains imaged by seismic tomography that revealed parallel belts of mid-mantle slabs below western NAM and the eastern Pacific (Clennett et al., 2020; Fuston & Wu, 2021; Sigloch & Mihalynuk, 2013, 2017; Sigloch et al., 2008; van der Meer et al., 2010, 2012, 2018). Below, we summarize the main geologic features of each terrane and key moments in their tectonic evolution.

The Alexander terrane is interpreted as a microcontinent that was intruded by late Neoproterozoic to early Paleozoic arc-related rocks (~600–400 Ma; Gehrels & Saleeby, 1987; White et al., 2016). Volcanic and plutonic rocks dominate the southern part of the Alexander terrane, whereas its northern part consists mainly of Paleozoic shelf strata as old as the late Cambrian (Beranek et al., 2012). Magmatic activity in the Alexander terrane was interrupted during two orogenic events, the Cambrian Wales orogeny and the Early Devonian Klakas orogeny (Gehrels & Saleeby, 1987).

The Wrangellia terrane consists mainly of Late Devonian to Early Permian and Late Triassic to middle Cretaceous arc-related igneous and sedimentary rocks (Alberts et al., 2021; Nokleberg et al., 1994; Plafker & Berg, 1994). Although basement rocks underlying the Paleozoic arc are not exposed, detrital zircon data from the southern Wrangellia terrane suggest that it incorporated fragments of the Alexander terrane (Alberts et al., 2021). Geologic ties between the Alexander and Wrangellia terranes dating back as early as ~360 Ma (Late Devonian) are corroborated by coeval gabbro complexes that intruded both terranes (Israel et al., 2014).

An episode of plume-related volcanism occurred between the Paleozoic and Mesozoic arc phases of the Wrangellia terrane. The up to 6 km-thick Wrangellia large igneous province (LIP) (Late Triassic, ~232–226 Ma; Greene et al., 2010) is preserved along the full length of the Wrangellia terrane. In our study area on Vancouver Island, the Bonanza arc was built on these flood basalts and associated sedimentary rocks (Canil et al., 2010, 2013; DeBari et al., 1999; D'Souza et al., 2016). The Bonanza Group consists of a basal Upper Triassic–lowermost Jurassic volcanic-sedimentary succession (Parson Bay Formation and Volcaniclastic-Sedimentary unit; ~226–200 Ma, based on detrital zircon U-Pb geochronology and biostratigraphy) overlain by the Lower Jurassic Le Mare Lake Volcanic Unit (~201–190 Ma, based on zircon U-Pb geochronology and biostratigraphy) and the Lower to Middle Jurassic Holberg Volcanic Unit (~201–164 Ma, based on detrital zircon U-Pb and amphibole Ar-Ar geochronology; Nixon & Orr, 2007; Nixon, Hammack, Hamilton, et al., 2011; Nixon,

Hammack, Koyanagi, Payie, et al., 2011; Nixon, Hammack, Koyanagi, Snyder, et al., 2011; Nixon, Snyder, et al., 2011). The younger volcanic units are interbedded with minor marine and non-marine epiclastic rocks and limestone. The bulk of the Bonanza arc complex was intruded by the Early to Middle Jurassic Island Plutonic Suite (~201–164 Ma, based on zircon U-Pb and amphibole Ar-Ar geochronology; D'Souza et al., 2016; Nixon, Hammack, Hamilton, et al., 2011, Nixon, Hammack, Koyanagi, Payie, et al., 2011; Nixon, Hammack, Koyanagi, Snyder, et al., 2011; Nixon, Snyder, et al., 2011; Nixon, Snyder, et al., 2011). The Le Mare Lake Volcanic Unit and the Island Plutonic Suite represent the main phase of growth of the Bonanza arc. On Vancouver Island, Triassic and younger rocks have experienced zeolite and prehnite-pumpellyite facies metamorphism, implying temperatures below 350°C and pressures below 2.5 kbar (<10 km depth; Kuniyoshi & Liou, 1976; Lei et al., 2020; Morris & Canil, 2022; Stewart & Page, 1974).

Except for the absence of flood basalts, the stratigraphic architecture of the Peninsular terrane is similar to that of the Wrangellia terrane (Figure 1). The Peninsular terrane mainly consists of Late Triassic to Middle Jurassic volcanic, plutonic, and volcaniclastic rocks of the Talkeetna arc, overlain by Upper Jurassic to Lower Cretaceous volcaniclastic basinal strata (McClelland et al., 1992; Plafker et al., 1989; Rioux et al., 2007, 2010). The Talkeetna arc, considered as an archetypal intraoceanic arc, is lithologically, temporally, and geochemically correlative to the Bonanza arc of the Wrangellia terrane (D'Souza et al., 2016; Plafker et al., 1989; Rioux et al., 2007). Scarce outcrops of metamorphosed upper Paleozoic mafic to intermediate volcanic rocks, limestones, and quartz-rich sedimentary rocks provide a glimpse into the basement of the Talkeetna arc (Plafker et al., 1989). Zircon xenocrysts from the Talkeetna arc and younger volcanic products suggest ties with the Alexander and Wrangellia terranes from at least the early Carboniferous (~310 Ma; Amato et al., 2007; Bacon et al., 2012; Beranek et al., 2014).

Based on stratigraphic, structural, geochemical, and geochronologic data from arc and accretionary complex rocks, the polarity of the subduction beneath the Wrangellia superterrane is thought to have been northward in western Alaska and eastward in British Columbia (in present-day coordinates) at least since ~200 Ma (Amato et al., 2013; Clift, Draut, et al., 2005, Clift, Pavlis, et al., 2005; Plafker & Berg, 1994; Trop & Ridgway, 2007). Notably, blueschist-facies metamorphic rocks were formed during the Early Jurassic in a subduction complex now juxtaposed against the outboard (seaward) margin of the superterrane (Roeske et al., 1989; Sisson & Onstott, 1986).

The timing, latitude, and possible diachroneity of the Wrangellia superterrane's accretion to NAM, are not yet resolved. Basinal records in the suture zone between the northern Wrangellia superterrane and the Intermontane superterrane to the east, have been used to suggest timings of collision that include the Middle Jurassic (e.g., McClelland et al., 1992), the Late Jurassic (e.g., Trop & Ridgway, 2007), the Early Cretaceous (e.g., Hampton et al., 2010), the middle Cretaceous (e.g., Amato et al., 2013; Plafker et al., 1989), and the Late Cretaceous (e.g., Hults et al., 2013). Coeval magmatism and development of thrust belts in the Coast Plutonic Complex are inferred to have initiated shortly after collision along the Wrangellia–Intermontane boundary, at ~100 Ma (Journeay & Friedman, 1993; Monger et al., 1982; Rusmore & Woodsworth, 1991). Proponents of diachronous accretion suggest that the southern part of the superterrane in British Columbia collided during the Jurassic, whereas the northern part of the superterrane in Alaska collided during the Late Cretaceous (e.g., Manselle et al., 2020; Trop & Ridgway, 2007). These accretion scenarios mostly differ from each other in the amount of post-subduction translation that the Wrangellia superterrane experienced, based on paleomagnetic studies (Kent & Irving, 2010).

Notably, estimates of post-accretion translation of the Wrangellia superterrane based on detrital zircon data from northern Washington, southern British Columbia, and southern Alaska have not reached any consensus, with offsets ranging from ~500 km (Mahoney et al., 1999, 2021), through ~1,000 km (Yokelson et al., 2015), to >1,500 km (Boivin et al., 2022; Housen & Beck, 1999; Matthews et al., 2017; Sauer et al., 2019). Among these studies, the main disagreement revolves around the source of detrital zircons from the Late Cretaceous Nanaimo Group of Vancouver Island that was deposited on the Wrangellia superterrane, which may have been sourced from rocks outcropping in Idaho (~300–600 km south from Vancouver Island) and/or southern California (~1,500–1,800 km south from Vancouver Island; Boivin et al., 2022, and references therein). In recent models of Late Cretaceous tectonic evolution of the Wrangellia superterrane (e.g., Hildebrand, 2013; Johnston, 2008; Tikoff et al., 2023), both options are compatible with paleomagnetic data from the Late Cretaceous Nanaimo Group,

which have consistently yielded paleolatitudes ~1,600–2,500 km south of its present-day location (Enkin et al., 2001; Kim & Kodama, 2004; Krijgsman & Tauxe, 2006; Ward et al., 1997).

#### 2.2. Jurassic Ophiolites and Accreted OPS of California

Westernmost California and southern Oregon display a Mesozoic–Cenozoic subduction complex that formed by episodic accretion of OPS, which records the history of a subducting oceanic plate from its formation to arrival at the trench (Isozaki et al., 1990). This subduction complex–the Franciscan Complex–accreted below suprasubduction zone ophiolites of Middle to Upper Jurassic age (~170–155 Ma) that are preserved as isolated klippen (Wakabayashi, 2015; Figure 1). The oldest accreted rocks are high-temperature, high-pressure metabasites interpreted as metamorphic sole rocks with Lu/Hf garnet ages of 180 Ma that mark the (minimum) age of subduction initiation (Mulcahy et al., 2018) and show that the ophiolites formed above an active subduction zone (e.g., Guilmette et al., 2018), consistent with their geochemical composition (Snortum & Day, 2020). Using GPlates combined with paleomagnetic data, Arkula et al. (2023) reconstructed western North American deformation to restore the relative positions of the Californian ophiolites in the Jurassic. Their paleomagnetic results also imply that the forearc paleo-ridges that generated the Jurassic ophiolites had near-perpendicular orientations to that of the Franciscan subduction zone. The kinematic restoration of the ophiolite belt shows that the paleoridges may have accommodated spreading rates of  $\sim 6$  cm/yr, suggesting that the plate subducting obliquely beneath California in the Jurassic had a northward motion relative to NAM of up to  $\sim 6-7$  cm/yr.

The Franciscan Complex consists of rocks accreted from oceanic crust that formed since the Early Jurassic (Wakabayashi, 2015). The youngest accretion may have happened as recently as 12 Ma (McLaughlin et al., 1982). Among the accreted rock assemblages, two localities with middle Cretaceous limestone blocks provided paleomagnetic results that allow computing a paleolatitudinal journey of the OPS prior to accretion. The Laytonville limestone (103–90 Ma) yielded paleolatitudes of  $17^{\circ} \pm 7^{\circ}$  (Alvarez et al., 1980) and  $14^{\circ} \pm 5^{\circ}$  (Tarduno et al., 1986). The Calera limestone (129–90 Ma), located ~300 km southeast of Laytonville, yielded paleolatitudes of  $24^\circ \pm 4^\circ$  (Courtillot et al., 1985) and  $18-25^\circ$  at 105–90 Ma (Tarduno et al., 1985). Because the blocks may have rotated during accretion to the Franciscan Complex, it is not obvious whether they formed on the southern or northern hemisphere. Also, the accretion age of the limestone blocks to the Franciscan Complex has uncertainties. All scenarios show that in Early to Late Cretaceous times, the OPS sequences that accreted to the Franciscan Complex were derived from lower latitudes, moving north at rates varying from ~8 cm/yr for the northern hemisphere options (Courtillot et al., 1985; Tarduno et al., 1985) to  $\geq 15$  cm/yr for the southern hemisphere options (Alvarez et al., 1980; Tarduno et al., 1986). Additionally, ocean plate igneous and sedimentary rocks with ages ranging from 180 to 110 Ma accreted to the Chortis block around 100 Ma (Andjić et al., 2019) at a latitude of 11°N (Costa Rica): those rocks yielded paleolatitudes of 8–20° north or south (Boschman et al., 2021).

## 3. Paleomagnetic Sampling and Measurements

To test the robustness of the paleomagnetic data from pre-accretionary rocks of the Wrangellia superterrane, we collected a total of 295 samples for paleomagnetic analysis from four localities in the Bonanza Group of northern Vancouver Island, Canada (Figure 2; Nixon, Hammack, Hamilton, et al., 2011; Nixon, Snyder, et al., 2011). Locality YM ("Yreka Main") was sampled in a continuous ~50 m monoclinal succession of impure limestones of the Norian–Rhaetian Parson Bay Formation, from which 119 samples were collected (average dip direction = 230°, average dip = 50°, n = 8). Basaltic to andesitic lava flows of the Hettangian–Sinemurian Le Mare Lake Volcanic Unit were sampled in three sections (MD = "Main Drive," TT = "Teeta Creek," VL = "Victoria Lake") that are ~5 km apart, with each section consisting of ~15 m (TT: average dip direction = 226°, average dip = 38°, n = 3) to ~50 m (MD: average dip direction = 273°, average dip = 29°, n = 14; VL: average dip direction = 196°, average dip = 33°, n = 12) of lavas. In our paleomagnetic analysis, we include data from 13 sites sampled by Irving and Yole (1987, abbreviated as I&Y87) along the Cape Perkins locality (average dip direction = 294°, average dip = 52°, n = 4) which is located ~40 km WNW from our sampling localities.

We used a gasoline-powered motor drill to sample 2.5 cm-diameter paleomagnetic cores, the orientation of which was measured with a magnetic compass with an inclinometer attached. We followed procedures recommended by Gerritsen et al. (2022) and drilled one core per limestone bed or volcanic flow (Figure S1 in Supporting Information S1) to optimize the amount of individual spot readings of the magnetic field, and as a field test selected a





**Figure 2.** Geological map of the studied area in northern Vancouver Island (100 m contour interval; modified after Nixon, Hammack, Hamilton, et al. (2011); Nixon, Snyder, et al. (2011)), which is shown in the lower left inset. The red stars on the geological map indicate the four localities where paleomagnetic samples were collected. Localities VL, TT, and MD are in the Le Mare Lake Volcanic Unit and locality YM is in the Parson Bay Formation. Parts of the Parson Bay Formation with higher volcaniclastic contents are indicated with a sandy lithological pattern.

total of seven lava sites where 5 cores per lava flow were drilled to evaluate whether within-site scatter of paleomagnetic data is low (i.e., k values typically exceeding 50; e.g., Johnson et al., 2008). Measurements were corrected for the local declination (16°47′E to 16°48′E).

The cores were processed at the Paleomagnetic Laboratory Fort Hoofddijk at Utrecht University, The Netherlands. The cores were cut into 2.2 cm-long samples using a double-blade circular saw. To determine the nature of magnetic carriers for both types of sampled lithology (lavas and impure limestones), thermomagnetic analyses were performed using a horizontal translation-type Curie balance with a sinusoidally cycling applied magnetic field, usually 100-300 mT (Mullender et al., 1993). Several heating-cooling cycles were applied to detect magneto-mineralogical alterations during heating. We used the following temperature scheme (in °C): 150, 75, 225, 150, 300, 225, 375, 300, 450, 375, 525, 450, 600, 20 (for lavas); 250, 150, 350, 250, 450, 350, 520, 420, 620, 500, 700, 20 (for limestones). Stepwise thermal (TH) demagnetization was applied to 117 limestone samples and 30 lava samples, whereas stepwise alternating field (AF) demagnetization was applied to 72 limestone samples and 174 lava samples, the latter processed with a robotized magnetometer (Mullender et al., 2016). Natural remanent magnetizations (NRM) were measured on a 2G DC SQUID magnetometer. Temperature steps of 100, 180, 210, 240, 270, 300, 320, 340, 360, 380, and 400°C were used for TH treatment of 87 limestone samples. Temperature steps of 100, 180, 210, 240, 270, 300, 330, 360, 390, 420, 450, 480, and 510°C were used for TH treatment of 30 other limestone samples. Temperature steps of 100, 180, 210, 240, 270, 300, 330, 360, 390, 420, 450, 480, 510, 540, 570, and 600°C were used for TH treatment of the 30 lava samples. Demagnetization steps of 5, 10, 15, 20, 25, 30, 35, 40, 45, 50, 60, 70, 80, 90, 100, and 120 mT were used for AF treatment of all samples. To improve the resolution of AF demagnetization results of limestones, the latter were heated to 150°C in a thermal demagnetizer prior to AF treatments (van Velzen & Zijderveld, 1995).

Sample interpretation and statistical analysis were conducted using the online portal Paleomagnetism.org (Koymans et al., 2016, 2020). All results can be imported into the portal from data files (.col) available in Data Sets S1 and S2,

as well as in the Paleomagnetism.org 2.4.0 database (Koymans et al., 2020). Demagnetization diagrams were plotted on orthogonal vector diagrams (Zijderveld, 1967), and the magnetic components were determined through principal component analysis (Kirschvink, 1980). Remanence directions in orthogonal vector diagrams were determined from at least four successive demagnetization steps, as follows: (a) principal component fits were anchored when demagnetization segments trend clearly toward the origin; (b) principal component fits were not anchored when demagnetization segments do not trend toward the origin. Great circle solutions were determined using the method of McFadden and McElhinny (1988). The fold test (Tauxe & Watson, 1994) and the bootstrap common mean direction test (Heslop et al., 2023) were used when applicable. The elongation-inclination (E/I) method (Pierce et al., 2022; Tauxe & Kent, 2004; Tauxe et al., 2008) was applied to the sedimentary locality. A maximum angular deviation cut off (i.e., MAD  $\leq 15^{\circ}$ ) was not applied to our data set because it only reduces the number of samples, which decreases paleopole precision (Gerritsen et al., 2022). Paleomagnetic pole positions were calculated using Fisher (1953) statistics on virtual geomagnetic poles (VGPs)-whereby each VGP is derived from a single site-following statistical procedures described in Deenen et al. (2011), providing a measure of the VGP dispersion (Fisher precision parameter, K) and a 95% confidence ellipse on the pole position (A<sub>95</sub>). The mean paleomagnetic direction and the 95% confidence regions on the declination ( $\Delta D_{y}$ ) and inclination ( $\Delta I_{y}$ ) were computed from the paleomagnetic pole and its A<sub>95</sub>. A 45° cutoff was applied to the VGPs (Johnson et al., 2008) at the group level for the volcanic localities (TT, MD, VL). Application of this cutoff to the sedimentary data set (YM) does not lead to the exclusion of any of the directions.





**Figure 3.** Rock magnetic results from the Parson Bay Formation (YM) and Le Mare Lake Volcanic Unit (TT, MD, and VL). (a), (c), (e), (g) Thermomagnetic curves measured on a Curie balance. Heating segments are in red and cooling segments are in blue. (b), (d), (f), (h) Intensity decay curves measured during thermal demagnetization.

#### 4. Results

#### 4.1. Volcanic Rocks of the Le Mare Volcanic Unit

Thermomagnetic curves show that lava samples have Curie temperatures close to  $580^{\circ}$ C (Figures 3g and 3h), indicating that magnetite is the main magnetic carrier. In orthogonal vector plots, the majority of the VL (58/80) and MD (32/54) samples reveal an overprint at low temperature/coercivity steps (up to ~100–240°C or 5–15 mT).





**Figure 4.** Orthogonal vector diagrams of representative samples from the Parson Bay Formation (YM) and Le Mare Lake Volcanic Unit (TT, MD, and VL) in geographic (= in situ) coordinates, where closed (open) symbols indicate declination (inclination). (a) Six samples from lava flows, including two opposite-polarity directions. (b) Five samples from limestones, including two opposite-polarity directions. ChRM = characteristic remanent magnetization.

VL samples (46/80) have an additional overprint at mid to high-range temperature/coercivity steps (~420–570°C or 15–90 mT), most of which (40/46) did not yield characteristic remanent magnetization (ChRM) directions. In contrast, TT samples show a linear decay toward the origin of demagnetization diagrams without overprints. Initial intensities ranged 7–550 mA/m for TT, 0.2–38 mA/m for MD, and 0.6–1,650 mA/m for VL. ChRM values were generally interpreted between 420 and 570°C or 10–120 mT for TT, ~270–510°C or ~25–70 mT for MD, and ~240–510°C or ~20–70 mT for VL. Interpreted ChRM directions from most samples yielded: (a) southward declinations and steep upward inclinations in in situ coordinates (Figures 4a and 5a); and (b) eastward declinations with shallow upward inclinations in tilt-corrected coordinates (Figure 5b). One TT sample and three MD samples yielded westward, down directions that suggest opposite polarity (Figure 4a); which of these two directions represent normal or reversed is not a priori known, as it depends on the hemisphere of origin. In the MD section, reversed magnetic polarities are located at the level of samples MD 1.23, MD 1.24, and MD 1.40; opposite-polarity ChRM directions of locality MD share a common mean direction.





**Figure 5.** Paleomagnetic results from the Parson Bay Formation (YM) and Le Mare Lake Volcanic Unit (TT, MD, and VL). Data from Irving and Yole (1987) are included. (a), (b), (f), (g) ChRM directions from localities TT (in red), MD (in blue), VL (in green), I&Y87 (in black), and YM (in gray). (a), (f) In situ coordinates. (b), (g) Tilt-corrected coordinates. (c), (d), (h), (i) Mean directions, including confidence intervals (dashed lines), of localities TT, MD, VL, I&Y87 and YM. (e) Mean direction of all volcanic localities, including confidence intervals (dashed lines). is = in situ coordinates, tc = tilt-corrected coordinates. (j)–(m) Bootstrapped fold tests of the sampled localities, with cumulative distribution function (confidence interval in light gray) based on 1,000 bootstrap samples (mean shown in red).

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Irving and Yole (1987) sampled two localities—Cape Perkins and Victoria Lake-in the Le Mare Lake Volcanic Unit of northern Vancouver Island. Their sites 5 and 8 were sampled along the same section as our locality VL. These two sites have northward declination and downward inclination values  $(D/I = ~025^{\circ}/~40^{\circ})$  that are incompatible with the average eastward declination and upward inclination  $(D/I = \sim 063^{\circ}/\sim 58^{\circ})$  of our data from the same locality. Directions from their sites 5 and 8 are actually comparable to overprints interpreted at mid- to high-range temperature/coercivity steps from the majority of our VL samples (46 out of 80), many of which have very shallow downward inclinations and northward declinations, and did not yield any interpretable ChRM directions (e.g., samples VL1.22–1.25, VL 1.43, VL 1.58–1.70; Data Set S1). Our sample VL1.36A shows a typical example where both a ChRM and an overprint have been interpreted (Figure 4a). In contrast, 13 sites from Cape Perkins reported by Irving and Yole (1987) share a common mean direction (in tilt-corrected coordinates) with data from our locality TT (Figure 5i).

Fold tests per section are inconclusive. Differences in bedding orientation among VL, TT, MD, and I&Y87 localities were used in several fold tests, of which the following are considered positive: (a) MD combined with I&Y87 (confidence interval between 62% and 111%, Figure 5j); (b) TT combined with I&Y87 (confidence interval between 75% and 111%, Figure 5k). When the four volcanic localities are combined with the sedimentary locality YM, the fold test yields a result with a confidence interval of 80%–95% (Figure 5m), whereas the combination of the four volcanic localities yields a result with a confidence interval of 64%-85%. We consider that the clustering (<100%) in these regional fold tests may be the result of: (a) uncertainties in bedding tilt correction, which may be primarily related to lava flow emplacement on preexisting topography (e.g., Irving & Yole, 1987); (b) differential block rotations among the lava flow sections. We used the inclination-only method of Arason and Levi (2010) to compute the mean inclination of the four volcanic localities, which yields a mean  $(-50.7^{\circ} \pm 2.5^{\circ})$  that is comparable to the Fisher mean presented below  $(-49.4^{\circ} \pm 3.0^{\circ})$ . This implies that local rotations do not have a significant effect on our results from the Le Mare Lake Volcanic Unit.

The lava sites from which we collected multiple samples per site yielded low dispersion (e.g., k = 395.9 for TT 1.31-1.35 and k = 66.9 for MD 1.16-1.20; Figure S2 in Supporting Information S1), confirming that lava sites may be treated as spot readings of the field. The paleomagnetic pole computed from the lava sites is located at latitude =  $30.8^{\circ}$ S, longitude =  $339.1^{\circ}$ E ( $N = 132, K = 22, A_{95} = 2.7^{\circ}$ ), based on a mean paleomagnetic direction of Declination (D)  $\pm \Delta D_x = 102.2^{\circ} \pm 3.1^{\circ}$ , Inclination (I)  $\pm \Delta I_x = -49.4^{\circ} \pm 3.0^{\circ}$  (Figure 5e, Table 1). The D and I values of the combined volcanic localities differ significantly from the recent GAD field ( $D/I = 000^{\circ}/70^{\circ}$ ). The  $A_{95}$  value ( $A_{95min} = 1.7^{\circ} < A_{95} = 2.7^{\circ} < A_{95max} = 3.8^{\circ}$ ) satisfies the *N*-dependent reliability envelope of Deenen et al. (2011), suggesting that the observed VGP scatter can be explained by paleosecular variation (PSV).

#### 4.2. Sedimentary Rocks of the Parson Bay Formation

Thermomagnetic curves of the limestones are consistent with three types of magnetic carriers. Samples YM 1.91 and YM 1.118 have very low initial magnetization, even diamagnetic susceptibility, that is, the signal is dominated by the quartz glass of the sample holder. The curves are completely reversible after cycling to 250°C and show minor decay after the 350°C cycle, with a pyrite-to-magnetite oxidation signature displayed after 420°C (e.g., Passier et al., 2001). On cooling to room temperature after the 700°C cycle, a discontinuity at  $\sim$ 320°C represents the Curie temperature of pyrrhotite (Figure 3c and Figure S2 in Supporting Information S1), which is produced during the thermomagnetic experiment. Pyrrhotite is stable with respect to pyrite at high temperatures and apparently not all pyrite is oxidized. If originally present, pyrrhotite would have shown up in the first heating segments as well. In contrast, samples YM 1.52 and YM 1.61 present a much higher (~50-100 times) initial magnetization, with smooth curves that are essentially reversible up to 500°C, which supports the presence of titanomagnetite with a variable Ti content (Figure 3e and Figure S2 in Supporting Information S1). Interestingly, both groups of samples have similar values of their initial NRM (YM 1.52 = 3.3 mA/m, YM 1.61 = 4.1 mA/m vs. YM 1.91 = 6.0 mA/m, YM 1.118 = 5.1 mA/m), which suggests that the efficiency of the NRM acquisition mechanism differs widely between the two groups. YM 1.91 and YM 1.118 may feature a chemical remanent magnetization, which is an efficient NRM acquisition mechanism. The NRM acquisition of YM 1.52 and YM 1.61 is much less efficient; a lot more magnetic material is required for a similar NRM intensity, which may imply a detrital remanent magnetization. Nonetheless, after thermal demagnetization at 360°C, usually <10%-15% of the initial NRM intensity remains for both types of samples. This result implies that ChRM directions were interpreted from the same temperature steps for both groups. Finally, sample YM 1.2 shows a largely reversible

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Table 1																						
Paleomag	netic Resu	Its of the Le 1	Mare L	ake Volca	nic Unit (MI	D, VL, Ti	$\Gamma$ ) and $ti$	he Parso.	n Bay .	Formatic	on (YM)											
							In si	tu			Tilt-corr	ected	ĺ									
Locality	slat	Slon	N	N45 (is)	N45 (tc)	D	$\Delta D_x$	Ι	$\Delta I_x$	D	$\Delta D_x$	Ι	$\Delta I_x$	k	$\alpha_{95}$	K	A <sub>95Min</sub>	$A_{95}$	$\mathrm{A}_{95\mathrm{Max}}$	r	plat	plong
Le Mare L	ake Volca	nic Unit																				
MD	50.357	-127.439	4	n.a.	n.a.	143.9	15.3	-72.0	5.3	120.3	7.5	-48.9	7.4	15.0	5.8	12.0	2.6	6.5	7.6	-29.8	-43.0	323.6
٨L	50.390	-127.392	29	n.a.	n.a.	144.2	37.4	-81.7	5.2	62.6	10.8	-58.2	7.7	18.9	6.3	11.2	3.1	8.4	9.8	-38.8	-16.2	7.3
TT	50.376	-127.546	59	n.a.	n.a.	161.0	6.2	-73.7	1.9	6.66	2.5	-48.6	2.6	98.7	1.9	71.3	2.4	2.2	6.3	-29.6	-28.6	335.6
I&Y87	50.450	-128.050	13	n.a.	n.a.	5.3	28.6	-76.2	7.2	96.1	6.5	-42.2	7.9	49.7	5.9	49.4	4.3	6.0	16.3	-24.4	-22.9	333.7
Mean	50.380	-127.528	145	126	132	153.3	6.4	-74.4	1.9	102.2	3.1	-49.4	3.0	28.8	2.3	22.0	1.7	2.7	3.8	-30.3	-30.8	339.1
Parson Ba	y Formatic	u																				
ΥМ	50.415	-127.512	139	n.a.	n.a.	347.5	4.1	65.8	2.0	279.2	2.0	40.3	2.5	38.0	2.0	44.7	1.7	1.8	3.7	23.0	23.7	151.3
<i>Note</i> . Data cutoff in it cone of cc A <sub>95min</sub> = r applicable.	from Irvir. 1 situ coord nfidence a naximum 8	ig and Yole ( tinates (is) an round the co and minimur	1987) a d after mputec 1 value	re includec tilt correcti 1 site mean of A <sub>95</sub> exF	d. $slat/slon =$ ion (tc); $D, \Delta$ i direction; H pected from J	I atitude $D_x = dec$ $\zeta$ and $A_g$ paleosecu	and long clinatior. <sub>55</sub> = pre- 11ar vari	gitude of and asso cision pa ation of	sampli ociated tramete the gec	ng locati error; <i>I</i> , r and se magneti	ion; $N = \Delta I_x = In$ $\Delta I_x = In$ miangle c field; 1	total nun Iclination of the 9. I = paleo	nber of 1 and as 5% cor 1 atitud	demagr sociatec le of co	netized 1 error; nfidenc localit	samples $k$ and $\alpha$ be aroun $\alpha$ ; $\gamma$ ; $\gamma$ plat/p	3; N45 = 95 = precent of the model of the	numbe cision p ean virt aleopole	r of samp arameter tual geom e latitude	les that fa and semia lagnetic F and long	ll within t ıngle of th ole; A <sub>95n</sub> ítude; n.a.	he 45° ie 95% <sub>ax</sub> and = not



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thermomagnetic curve with a Curie temperature close to 320°C (Figure 3a), suggesting that an iron sulfide, likely pyrrhotite, is present in the limestone section.

Initial NRM intensities ranged 0.3–400 mA/m, with values mostly restricted to 2–6 mA/m for samples from which ChRMs were interpreted. In orthogonal vector plots, magnetic components could not be determined from a subset of samples (48/189) because of erratic demagnetization behavior. The maximum applicable AF (120 mT) was not high enough to fully demagnetize limestone samples. Nevertheless, ChRM directions interpreted from AF demagnetization data agree well with those obtained from TH demagnetization data, as they share a common mean direction. A few samples (5/140) show an overprint at low temperature/coercivity steps (~100–210°C or 5–20 mT). ChRM values were generally interpreted at ~210–360°C or ~30–70 mT, yielding down, westward directions. Seven YM samples yielded eastward, up directions that suggest opposite polarity; reversed magnetic polarities are located at the base of the section at the level of samples YM 1.1A, YM 1.2A (Figure 4b), YM 1.19A, YM 1.21A, YM 1.22A (Figure 4b), YM 1.22B, and YM 1.23A. Opposite-polarity ChRM directions of locality YM share a common mean direction. The paleomagnetic pole computed for locality YM is located at latitude = 23.7°N, longitude = 151.3°E (N = 139, K = 44.7,  $A_{95} = 1.8°$ ), providing a direction with  $D \pm \Delta D_x = 279.2° \pm 2.0°$ ,  $I \pm \Delta I_x = 40.3° \pm 2.5°$  (Figure 5g, Table 1). The *D* and *I* values of YM in tilt-corrected coordinates differ significantly from that of the recent GAD field (D/I = 000°/70°).

Overall, we find that the results of the YM limestones are consistent with dominantly carrying a chemical remanent magnetization. This remanence was acquired soon after sediment deposition and resulted in ChRM acquisition that possibly integrated secular variation over a duration of  $\geq 10^3$  yr. This interpretation is compatible with the following aspects: (a) All ChRMs are interpreted between 210 and 360°C and present very high coercivities (no full demagnetization at 120 mT), suggesting that pyrrhotite, a diagenetic mineral, could be the main magnetic carrier of YM samples (e.g., Pastor-Galán et al., 2016). (b) The relatively low dispersion of the VGPs of the limestones yield an  $A_{95}$  (1.8°) that is slightly higher than the  $A_{95min}$  (1.7°) of the reliability envelope of Deenen et al. (2011); (c) We computed the elongation of the YM data set (in tectonic coordinates) to compare it with the predicted elongation versus inclination curve of the TK03.GAD model of Tauxe and Kent (2004), following the approach of Pierce et al. (2022; Figure S2 in Supporting Information S1). We find that the elongation (~1.9) predicted by the TK03.GAD curve falls within the bootstrapped 95% confidence range of the YM elongation (1.1-2.6). This indicates that the shape of the directional distribution is consistent with the paleosecular variation model, which is compatible with the acquisition of a chemical remanent magnetization during (early) diagenesis (Pierce et al., 2022). This interpretation is further supported by the fact that YM limestones (age range of about 227 to 201.3 Ma) share a common mean direction with slightly younger lava flows (our TT locality and Cape Perkins of Irving & Yole, 1987, p. 201.3-190.8 Ma; Figure 5i). Moreover, directions of YM limestones and those of Cape Perkins lavas of Irving and Yole (1987) pass a fold test (confidence interval between 98% and 125%; Figure 51).

Although our preferred interpretation is that of an early diagenetic magnetization, we cannot completely rule out that the YM limestones were remagnetized after the Late Triassic, especially because it was not possible to perform a fold test within the Parson Bay Fm. YM limestones could have been remagnetized at two times: (a) during the Earliest Jurassic, the emplacement of the Bonanza arc could have remagnetized the underlying limestones, which is compatible with the close agreement between YM and TT mean directions; (b) during the Cretaceous–Eocene, remagnetization could have occurred after the tilting of the limestones, which could explain the agreement between the YM mean direction in geographic coordinates and the expected declination ( $\sim$ 340°) and inclination ( $\sim$ 60°) for the sample location for this time period. However, we find the second possibility unlikely, because the geographic mean direction of the YM limestones is comparable to the northward declinations and steep, downward inclinations of the opposite-polarity, geographic mean directions of volcanic sections (TT, VL, and MD).

#### 4.3. Summary of the Results

There is no independent control on the hemispheric origin and direction of rotation (clockwise or counterclockwise) for the YM and combined MD–TT–VL localities, because both normal and reversed magnetic polarities are common in the Norian–Rhaetian and Hettangian–Sinemurian, respectively. Therefore, for each paleomagnetic data set, there are two possible solutions corresponding to either the northern or southern hemisphere, with opposite senses of rotation. The northern hemisphere option for the sediment-derived data set (YM)

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yields a paleolatitude of  $23.0^{\circ} \pm 2.5^{\circ}$  N and a ~81° counterclockwise rotation (relative to present-day spin axis), whereas the southern hemisphere option yields a paleolatitude of  $23.0^{\circ} \pm 2.5^{\circ}$  S and a ~99° clockwise rotation. The combined result of the data from igneous rocks of the Bonanza Group (MD-TT-VL-I&Y87) yields either a northern hemisphere solution with an estimated paleolatitude of  $30.3^{\circ} \pm 3.0^{\circ}$ N and a ~78° counterclockwise rotation, or a southern hemisphere solution with a paleolatitude of  $30.3^{\circ} \pm 3.0^{\circ}$ S and a ~102° clockwise rotation.

## 5. Updated Paleomagnetic Data Compilation of the Wrangellia Superterrane

We combine our new data with available paleomagnetic data from the Wrangellia superterrane that we compiled from the literature. We compiled paleomagnetic data from rocks that are located within the boundaries of the Wrangellia superterrane (Figure 1), which includes rocks that were formed before (e.g., Bonanza Group, Karmutsen Formation) and after the accretion of the superterrane (e.g., Nanaimo Group), during the middle Cretaceous. We only chose data from the NW-SE striking part of the Wrangellia superterrane, from southern British Columbia to eastern Alaska. The western and northern parts of Alaska (i.e., Peninsular terrane) were not included in the compilation, as these were proposed to have been rotated in Late Cretaceous to Paleocene times during oroclinal bending (Box, 1985; Dover, 1994; Johnston, 2001). Our data compilation includes data sets that contain at least eight individual directions of the magnetic field (either from eight individual cooling units in magmatic rocks, or eight sedimentary beds, as suggested by Meert et al., 2020), and of which the distribution of VGPs passes the Deenen et al. (2011) criterium of representing paleosecular variation ( $A_{95min} < A_{95} < A_{95max}$ ). We included sediment-derived data sets that were not corrected for potential inclination shallowing and thus do not satisfy the recently proposed reliability criteria for sedimentary data sets of Vaes et al. (2021). We do so to increase the temporal coverage of our data compilation, including, for instance, the data sets derived from Paleozoic (Bazard et al., 1995; Van der Voo et al., 1980) and Upper Cretaceous sedimentary rocks. We excluded data sets from the compilation if the authors of the original study interpreted the magnetic signal to represent a remagnetization (Butler et al., 1997; Haeussler, Coe, & Onstott, 1992; Hillhouse & Grommé, 1980; Irving & Massey, 1990; Symons, 1985), rocks with an unknown paleohorizontal were sampled (Irving et al., 1985; Irving & Massey, 1990; Rusmore et al., 2013; Symons, 1973), or shearing was interpreted to have influenced the magnetic directions (Butler et al., 2002).

Our compilation contains data sets for Early Jurassic and older times, and for Late Cretaceous and younger times. There are two small data sets from Lower Cretaceous strata in Alaska that were collected by Stone et al. (1982), and some preliminary results provided by Panuska et al. (1984). However, the original data and descriptions are not available, and the data sets were considered unreliable in subsequent studies (Butler et al., 1997; Harbert, 1990; Hillhouse, 1987). In addition, we excluded the paleomagnetic results from the mid-Cretaceous (~110 Ma) ultramafic rocks of Duke Island (Bogue & Grommé, 2004) due to the low number of sites (N = 6) and the uncertainties surrounding the determination of the paleohorizontal (Bogue & Grommé, 2004; Butler, Gehrels, & Saleeby, 2001). Butler et al. (1997) revisited Ordovician, Silurian, Devonian, and Carboniferous rocks of the Wrangellia superterrane from which Van der Voo et al. (1980) reported paleolatitudes. Butler et al. (1997) argued that the magnetizations carried by these Paleozoic rocks are largely secondary, concluding that none of these provide useful paleogeographic information. We show the data of Van der Voo et al. (1980) in our compilation but we consider them as potentially unreliable, as discussed further in Section 6.1.

The final compilation consists of a total of 38 collections from 17 studies, 23 for the Late Cretaceous and younger, and 15 for pre-Late Cretaceous times (Tables S1 and S2 in Data Set S3). For Late Cretaceous and younger times, paleolatitudes are all  $\sim 30^{\circ}$  or higher and interpreted as northern hemispheric. For the older data sets, paleolatitudes can be either southern or northern hemispheric. The exception are rocks from the Late Carboniferous-Permian Kiaman superchron ( $\sim 320-260$  Ma; Opdyke & Channell, 1996), which must be of reversed polarity if they carry a primary magnetization (Panuska and Stone, 1981). In Section 6.2, we show and discuss both hemispheric options.

#### 6. Discussion

#### 6.1. Paleomagnetic Data From the Wrangellia Superterrane: Paleogeographic Drift or Artifact?

To evaluate the paleomagnetically permissible plate motions of the Wrangellia superterrane since the Triassic, we use the updated paleomagnetic data compilation that includes our two new paleomagnetic poles. First, we briefly re-evaluate whether the new global APWP of Vaes et al. (2023) for the last 320 Ma provides significant



**Figure 6.** Declination (a) and paleolatitude (c) curves for the North American plate at a reference point at 50.4°N and 127.5°W, as predicted by three global APWPs and the high-resolution apparent polar wander path (APWP) for NAM by Gallo et al. (2023). The colored bands show the 95% confidence regions for each curve. Note that the global APWP of Kent and Irving (2010) and the APWP of Gallo et al. (2023) do not cover the entire time range discussed here. Markers show declination and paleolatitude values computed from paleomagnetic data sets from stable NAM (compiled by Vaes et al., 2023), with error bars that indicate the age uncertainty and 95% confidence in the declination or paleolatitude. Blue dots and white squares/diamonds indicate data sets that were accepted or rejected for inclusion in the global APWP of Vaes et al. (2023), respectively. The diamonds represent sediment-derived data sets that were not corrected for inclination shallowing, whereas the squares indicate data sets from igneous rocks. Relative rotations (b) and paleolatitudinal differences (d) between the accepted North American data and the global APWP are computed using the data-comparison approach of Vaes et al. (2022), in which the uncertainties of the reference poles provided by the APWP are weighted against the number of sites that underlie the investigated data sets. Discordant (i.e., statistically different from zero) results are colored in white with red edges and error bars.

modifications to the paleolatitude and declination reference curves for the North American continent that would alter the previous interpretations of Wrangellia-North America motion. This new APWP is based on (re-sampled) site-level paleomagnetic data rather than on study-mean paleomagnetic poles, which were previously used for APWP computation (e.g., Kent & Irving, 2010; Torsvik et al., 2012; see Vaes et al., 2022, 2023, for details). First, as illustrated in Figure 6, the global APWP of Vaes et al. (2023) gives a smaller uncertainty but is mostly within error of widely used global APWPs (in North American coordinates) of Kent and Irving (2010) and Torsvik et al. (2012). Second, we evaluate whether paleomagnetic data sets from stable NAM that contributed to the global APWP are anomalous. Therefore, we have plotted all data derived from NAM that were used in the construction of the global APWP of Vaes et al. (2023; Figure 6), as well as the North American data that were included in previous global compilations but did not satisfy the selection criteria used by Vaes et al. (2023). Most of the accepted North American data are in excellent agreement with the declination and paleolatitude predicted by the global APWP, including the Late Triassic and Early Jurassic data. When comparing these data directly with the predicted values of the global APWP using the data-comparison approach of Vaes et al. (2022), we find that only 3 out of 74 data sets ( $\sim$ 4%) have a statistically different declination, and 8 out of 74 data sets ( $\sim$ 11%) have a significantly different paleolatitude. These results are consistent with those obtained earlier by Enkin (2006), who found similar agreement between global and North American data. Third, we show the predicted declination and paleolatitude curves based on the 1-Ma-resolution North American APWP for the last 56 Ma of Gallo et al. (2023). These curves are again statistically indistinguishable from those of the global APWP of Vaes et al. (2023), although the curve of Gallo et al. (2023) predicts a paleolatitude that is  $\sim$ 3° lower between about 30

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and 56 Ma. Based on these observations, we regard the global APWP of Vaes et al. (2023) as a reliable reference curve against which the data from the Wrangellia superterrane can be compared.

Most previous studies assessed tectonic motions of mobile terranes such as the Wrangellia superterrane relative to NAM by using pole-to-pole comparisons between poles obtained from a mobile terrane and individual studymean poles of "cratonic" NAM (e.g., Dickinson & Butler, 1998; Irving & Wynne, 1990; Tikoff et al., 2023). However, the dispersion of reliable paleomagnetic reference poles from stable NAM (Figure 6) indicates that such an approach may be less robust compared to using a global APWP as a reference, which is based on considerably more data and has a significantly more robust temporal coverage (e.g., Enkin, 2006). This scatter among coeval paleomagnetic poles from stable parts of continents was previously observed by Harrison and Lindh (1982) and, more recently, Rowley (2019) and Vaes et al. (2022), and has been shown to be mostly caused by the undersampling of paleosecular variation of the past magnetic field. A considerable portion of the North American data that do not pass the selection criteria of Vaes et al. (2023) have been included in previous compilations of North American reference poles as well as for the computation of reference APWPs. The use of these poles may have influenced previous calculations of tectonic displacements of Cordilleran terranes relatively to the North American plate, particularly for Jurassic and older times. In addition, we note that different data selection criteria have led to substantially different Jurassic APWPs for NAM. For instance, the inclusion or exclusion of Jurassic paleomagnetic poles derived from Adria (rotated to North American coordinates; Muttoni & Kent, 2019) or from the Middle Jurassic Summerville and Morrison Formations (Mirzaei et al., 2021) have strong influence on the resulting APWP. We emphasize that the choice of a reference APWP will influence plate tectonic reconstructions of NAM and exotic terranes such as the Wrangellia superterrane.

Here, we compare the data of the Wrangellia superterrane with the global APWP in North American coordinates of Vaes et al. (2023) for the last 320 Ma, and with the running mean, 'pole-based' APWP for Laurentia of Torsvik et al. (2012) for earlier parts of the Paleozoic (Figure 7). Plotting the data compilation against a North American reference curve for both northern (N) and southern (S) hemisphere scenarios allows us to re-evaluate previous hypotheses on the primary or secondary nature of specific paleomagnetic data. First, both declinations and inclinations from Paleozoic to Early Jurassic rocks of the Wrangellia superterrane are systematically different from those of the Late Cretaceous to Cenozoic rocks. If a regional unrecognized remagnetization in Late Cretaceous or younger times had occurred, creating apparently low paleolatitudes by erroneously correcting for bedding tilt (Butler, Gehrels, & Kodama, 2001; Hollister et al., 2004; Housen & Beck, 1999; Monger & Price, 1996; Nelson & Colpron, 2007), then that remagnetization process would also have likely affected older rocks. The coherent, systematically deviating paleomagnetic directions of pre-Late Cretaceous rocks make a regional remagnetization in Late Cretaceous or younger times unlikely. Rather, remagnetization processes during those times are considered to be localized (e.g., Housen, 2018).

On the other hand, previous arguments for remagnetization of Paleozoic rocks (Butler et al., 1997) cannot be excluded. Rocks older than the Triassic have declinations as well as inclinations that are similar to those from the Triassic Wrangellia LIP (Figure 7). If they were not remagnetized during the eruption of the LIP, these rocks would suggest that for a period of ~250 Ma, from the Devonian to the Triassic, Wrangellia was part of a plate that was not moving much in paleolatitude, nor undergoing systematic vertical-axis rotations. We therefore avoid interpreting the pre-Triassic history of Wrangellia here in detail. However, the northern hemisphere-counterclockwise rotation scenario permits interpretation of the available Paleozoic data as primary. In contrast, the southern hemisphere-clockwise rotation scenario requires that at least the rocks from the Carboniferous-Permian Kiaman superchron do not carry a primary magnetization. Future detailed paleomagnetic study of these rocks may resolve this difference.

#### 6.2. Paleomagnetic Constraints on Wrangellia Plate Motion Since the Triassic

When taking the APWP of NAM as a reference, the paleomagnetic data set of the Wrangellia superterrane yields two hemispheric options with consistent declinations (Figure 7). These clusters imply either  $\sim 100^{\circ}$  clockwise or  $\sim 80^{\circ}$  counterclockwise rotation relative to the magnetic north pole since the Early Jurassic. Previous workers have mostly focused on paleolatitudes and assumed that declinations were unreliable to constrain the preaccretionary rotation of the Wrangellia superterrane because of local rotations related to post-accretionary orogenic deformation (e.g., Beck, 1976; Kent & Irving, 2010). Such local rotations could be explained, given that the Cordilleran orogen has been folded, faulted, and transported northward along the North American margin,





Scenario N - Multi-rotation/Northern hemisphere



Age (Ma)

possibly along major strike-slip faults (Beck, 1976, 1980; Irving & Yole, 1987). However, inspection of the data compilation does not suggest that local rotations play a major role. The scatter in declinations for collections of Late Cretaceous rocks of the Wrangellia superterrane is several tens of degrees. Much of these data collections are small, based on a dozen or so paleomagnetic sites (Tables S1 and S2 in Data Set S3). As mentioned above, even the declinations obtained from data sets of the stable plate interior of NAM (and provided input for global APWPs) are scattered over 30–40° (Figure 6a). This large dispersion of coeval paleomagnetic poles was previously shown to predominantly result from limited averaging of paleosecular variation (Harrison & Lindh, 1982; Vaes et al., 2022). The declination scatter of data from the Wrangellia superterrane is larger than that, suggesting that local tectonic rotations enhanced it. Nonetheless, the declinations of Triassic and Jurassic rocks are not chaotic: Lower Mesozoic and older rocks yield declinations that are systematically much larger than those of the

Age (Ma)

Upper Cretaceous rocks, but their scatter is similar to the latter. Similarly, the declination scatter of Paleozoic rocks from the Alexander and northern Wrangellia terranes is large but not chaotic. Importantly, declinations of Paleozoic rocks are strongly rotated and overlap those of Late Triassic rocks. We consider that these paleomagnetic data are not suitable to infer the paleogeographic evolution of the Alexander and northern Wrangellia terranes throughout the Paleozoic because they could have been the result of remagnetization during the Late Triassic. However, paleomagnetic data from Paleozoic rocks taken together with paleomagnetic data from Late Triassic–Early Jurassic rocks support the view that large magnitude rotations are the result of pre-accretionary plate motions of the Wrangellia superterrane, rather than the consequence of post-accretionary vertical-axis rotations linked to orogenic deformation. We therefore infer that the Wrangellia superterrane underwent a coherent rotation between the Early Jurassic and the Late Cretaceous, that is, during the pre-collisional period when it was part of a plate converging with NAM (Tikoff et al., 2023). In a southern hemisphere scenario (Scenario N) it represents a counterclockwise rotation of  $\sim 70^{\circ}$  (Figure 7).

Previous data from Triassic and older rocks of the Wrangellia superterrane yield paleolatitudes close to the equator (Figure 7). Our new data set shows a rapid paleolatitudinal motion on the order of  $20-30^{\circ}$  in the Late Triassic to Earliest Jurassic, which is northward in Scenario N, and southward in Scenario S (Figure 7). In Scenario S, this is followed by a northward motion of  $\sim 50^{\circ}$  in  $\sim 110$  Ma ( $\sim 5$  cm/yr). In Scenario N, paleolatitudes would remain fairly constant until the Late Cretaceous.

To interpret what such paleolatitudinal changes would mean for plate motion rates, we compare these trajectories with the paleolatitudinal motion of cratonic NAM. Scenario S gives a southward shift of the Wrangellia superterrane in the Triassic, followed by a northward motion (Figure 7). Interestingly, a southward trajectory is also visible from Late Triassic to Early Jurassic time in the global APWP in North American coordinates, superimposed on a northward trend. This 'dip' is observed worldwide and is interpreted as a major phase of True Polar Wander (TPW) that occurred in Late Triassic to Jurassic times (Steinberger & Torsvik, 2008; Torsvik et al., 2012; Vaes & van Hinsbergen, 2024). TPW is the tilting of the solid Earth relative to the spin axis, which results from changes in its moment of inertia, mostly because of the slow sinking of slabs through the mantle (Steinberger & Torsvik, 2010) and is directly detected as the difference between plate motions in a paleomagnetic reference frame (which includes TPW effects), and a mantle reference frame (which is blind to TPW effects; e.g., van Hinsbergen et al., 2015). It is important to realize for comparing Wrangellia's paleomagnetic data with a North American reference curve that the magnitude of TPW-induced paleolatitudinal change is paleolongitudedependent. Comparison between the paleomagnetic reference frame of Vaes et al. (2023) with the recent mantle reference frame of Müller et al. (2022) suggested that the pole of TPW was located at the equator (by definition), within the Atlantic Ocean (Vaes & van Hinsbergen, 2024), consistent with earlier estimates (Steinberger & Torsvik, 2008). At the longitude of western NAM, this TPW phase caused a southward and then northward shift in latitude on the order of 15° with a peak magnitude of TPW around 200 Ma. Farther west, in the eastern Panthalassa Ocean where the Wrangellia superterrane must have been, the magnitude of TPW increased up to a maximum of  $\sim 20^{\circ}$ . Much of the southward shift from the equator to  $> 20^{\circ}$ S in Scenario S could thus result from TPW rather than plate motion relative to the mantle. Because Jurassic motion of NAM was northward, northward motion of the Wrangellia superterrane relative to NAM was  $\sim 50^{\circ}$  between  $\sim 190$  and  $\sim 80$  Ma, a period of 110 Ma in Scenario S, that is, an average northward relative plate motion rate of  $\sim$ 5 cm/yr superimposed on a paleomagnetically unresolved eastward, convergent motion component (Figure 7). There are currently no highquality paleomagnetic data to further specify how this relative northward motion was distributed through time.

Scenario N requires a rapid northward shift of  $\sim 30^{\circ}$  of the Wrangellia superterrane in the Late Triassic to Earliest Jurassic. It is important to note that this motion is opposite to the TPW-induced southward motion, which means that the plate tectonic motion in Scenario N is larger, by about  $\sim 20^{\circ}$ , than the paleomagnetically determined motion. This would result in a relative northward plate motion component of Wrangellia of  $\sim 50^{\circ}$  between 220 and 190 Ma, that is, almost 20 cm/yr, and requires that it was followed by a net paleolatitudinal standstill of the plate carrying the Wrangellia superterrane until the Late Cretaceous. Because NAM continued moving northwards in the Jurassic, the plate carrying the Wrangellia superterrane would have to have moved southwards relative to NAM at a few cm/yr, that is, with a left-lateral strike-slip component. Below, we place both scenarios in further plate tectonic context.



# **Tectonics**



**Figure 8.** Schematic kinematic scenarios of the motion of the Wrangellia superterrane during the Late Triassic to Late Cretaceous (220–80 Ma). (a) Scenario N depicts a rapid Late Triassic–Early Jurassic northward shift of  $\sim$ 50° in the northern hemisphere, followed by a southward motion of lower amplitude, and then a standstill at middle latitude accompanied by counterclockwise rotation of  $\sim$ 70°. (b) Scenario S shows a Late Triassic–Early Jurassic southward motion of  $\sim$ 15° followed by a northward motion of  $\sim$ 5,000 km from the southern hemisphere accompanied by a clockwise rotation of  $\sim$ 110°. AFR, Africa; CAR, Caribbean; EUR, Europe; FAR, Farallon; IZA, Izanagi; IZG, Izanami; KUL, Kula; NAM, North America; PAC, Pacific; PHO, Phoenix; SAM, South America; WRA, Wrangellia.

#### 6.3. Wrangellia in Context of Eastern Panthalassa Plate Kinematic History

We explore the plate tectonic consequences and feasibility of the paleomagnetic scenarios above. A detailed kinematic restoration of Cordilleran orogenic architecture is required to fully justify a final choice between the two options. Such a reconstruction is beyond the scope of the current paper. We therefore restrict ourselves here to outlining the implications, solutions, and problems that the two scenarios generate.

We present a simplified tectonic reconstruction in Figure 8. We assume a 190 Ma-position of  $\sim$ 30° south or north of a simple, straight Wrangellia superterrane–which does not include the parts bent into Alaska (Johnston, 2001)– relative to NAM in the paleomagnetic reference frame of Vaes et al. (2023). Because we aim to develop a plate tectonic scenario and compare it to seismic tomographic constraints, we place the reconstruction in the TPW-corrected paleomagnetic reference frame of Torsvik et al. (2012). The latter frame has no paleolongitudinal constraints, so we restrict ourselves in the comparison to tomography and latitudinal fits only.

The paleolongitudinal position of the Wrangellia superterrane relative to the Americas is not constrained by paleomagnetism, but options are restricted given the estimated position of the Farallon-Phoenix-Izanagi triple junction at which the Pacific Plate formed at 190 Ma (Boschman & van Hinsbergen, 2016). The position of this triple junction must have been located in the eastern Panthalassa Ocean to maintain convergence of the major Panthalassa plates with the surrounding continents, and we use the approximated position of Boschman et al. (2021). This reconstruction gives the Wrangellia superterrane a possible paleolongitudinal range of  $\sim$ 5,000 km at 190 Ma.

From Late Triassic ( $\sim$ 210 Ma) to Middle Jurassic times ( $\sim$ 165 Ma), when the Bonanza arc was active, the Wrangellia superterrane was in an upper plate position of a subduction zone. In present-day coordinates, the trench is thought to have been located to the west of the Wrangellia superterrane (Clift, Pavlis, et al., 2005; Trop & Ridgway, 2007). This geometry requires that in Scenario S, the Wrangellia superterrane was underlain by a

northward dipping subduction zone. In Scenario N, the Wrangellia superterrane was underlain by a southward dipping subduction zone.

We evaluate the scenarios against the evidence for oblique subduction below the California forearc between 170 and 160 Ma with a N-S component of  $\sim 6-7$  cm/yr, deduced from reconstructing the Californian ophiolites (Arkula et al., 2023), and the evidence for a northward motion component throughout the Cretaceous as suggested by the paleolatitudes of the accreted seamounts of the Franciscan accretionary complex of California (Courtillot et al., 1985; Tarduno et al., 1985, 1986). In addition, a northward motion component is also permitted that is in line with the OPS that accreted to the western margin of the Chortis block in the Early Cretaceous, although a stable eastward relative plate motion is also permitted there (Boschman et al., 2021).

We use two additional constraints in the 190–80 Ma time window to illustrate the implications of the two scenarios. First, the Caribbean Plate contains Jurassic crust that formed around  $\sim$ 160–150 Ma, and that was part of the Farallon Plate prior to the  $\sim$ 100 Ma onset of subduction at the western Caribbean subduction zone (Boschman et al., 2019). This crust likely formed at the Farallon-Phoenix spreading ridge, and paleomagnetic data reveal that this crust was located near the equator during the Late Jurassic (Boschman et al., 2019). This requires that, by  $\sim$ 160 Ma, the Wrangellia superterrane must have been located to the north of the Caribbean lithosphere, and that it has likely always been located to the north of the Phoenix-Farallon spreading ridge.

Second, the Bonanza arc must have been associated with a subducting slab, which likely broke off around or shortly after the time of arc cessation, at  $\sim 165$  Ma. Subducting slabs may be dragged horizontally through the mantle over 1,000 km or more, driven by absolute plate motion of the subducting plate (Parsons et al., 2021; Qayyum et al., 2022; Spakman et al., 2018; van de Lagemaat et al., 2018), which implies that the present-day position of the slab remains of the Bonanza arc subduction zone may no longer reflect the location at which subduction started. However, tomography-plate reconstruction correlations suggest that slabs undergo no major horizontal motion after their break-off (Domeier et al., 2016; van der Meer et al., 2010, 2018), which suggests that the Bonanza arc-related slab is likely located in the lower mantle beneath the location where it broke off. This slab is likely one of the western belt slabs that have been identified below and west of NAM (Clennett et al., 2020; Sigloch & Mihalynuk, 2013, 2017; van der Meer et al., 2010, 2012). The southernmost of these, the Malpelo slab west of Colombia, was previously considered a candidate to be linked to the Bonanza arc (van der Meer et al., 2018). However, because the Caribbean and western South American Jurassic to Lower Cretaceous arcs have been located at that paleolatitude (Boschman et al., 2019), these arcs may provide a better candidate to explain the Malpelo slab. The southernmost slab of the western belt that cannot be explained by Caribbean arcs is the Socorro slab (van der Meer et al., 2010, 2018), at a latitude of  $\sim 17^{\circ}$ N (Figure 1). It is also not likely that this slab correlates to subduction records of the Guerrero terrane (in Mexico), which was built since Triassic time on accretionary prism rocks formed at the North American margin and was only temporarily separated from NAM by a short-lived and likely narrow back-arc basin (Boschman, Garza, et al., 2018; Boschman, van Hinsbergen, et al., 2018; Busby et al., 2023; Martini et al., 2014). Because the Bonanza arc is the southernmost intra-oceanic arc complex in the Cordilleran orogen, we therefore discuss in our scenarios S and N whether the Socorro slab may have been linked to the Bonanza arc.

Both scenarios share the post-80 Ma history in which paleomagnetic data place the Wrangellia superterrane about  $20^{\circ}$  south of its present latitude. The declination data are scattered and permit a rotated or a non-rotated position relative to today (Figure 7). In our schematic reconstructions, we place the Wrangellia superterrane in a non-rotated position, parallel to and close to the North American margin. We are well aware of the controversy around this reconstruction (e.g., Johnston, 2008; Trop & Ridgway, 2007), or even around the timing of the Wrangellia superterrane accretion (e.g., Gehrels et al., 2009; Hampton et al., 2010; McClelland et al., 1992; Monger, 2014; Nokleberg et al., 2000; Saleeby, 2000; Stevens Goddard et al., 2018; Tikoff et al., 2023; Trop et al., 2002), or diachroneity of accretion (e.g., Manselle et al., 2020; Nokleberg et al., 2000; Pavlis et al., 2019; Trop & Ridgway, 2007). At this stage, we have no satisfactory solution for where the inferred northward motion, which remains unaccounted for in the structural record, was accommodated. The position of the Wrangellia superterrane was located along the North American margin. However, the ~190–80 Ma history discussed below, during which the northern or southern hemisphere options play a role, would not change if we assumed a more northerly position for the Wrangellia superterrane at 80 Ma.

Scenario N requires ~5,000 km of northward motion of the Wrangellia superterrane relative to the mantle in the Late Triassic–Early Jurassic, which with southward subduction requires roll-back of the slab below the Bonanza arc. Between 190 and 80 Ma, the Wrangellia superterrane remained at middle latitudes until its collision with NAM, in which case the Bonanza arc was not associated with subduction of the Socorro slab. Another southward dipping subduction zone must therefore have existed between the Wrangellia superterrane and the Farallon Plate, which in Late Jurassic to Early Cretaceous time must have experienced a northward motion component (Arkula et al., 2023; Courtillot et al., 1985; Tarduno et al., 1985, 1986). Reconstructing the Wrangellia superterrane farther south around the mid-Jurassic, as depicted in Figure 8, would require slab advance of the Bonanza subduction zone, combined with another southward dipping subduction zone to accommodate the convergence between the Wrangellia superterrane and the Caribbean lithosphere that remained around equatorial latitudes (Boschman et al., 2019). Finally, the arrival of the Wrangellia superterrane at the North American margin must have involved a complex, counterclockwise plate rotation so that the western margin of the Wrangellia superterrane faced the Panthalassa Ocean. This counterclockwise rotation cannot be reconstructed without generating Farallon-Wrangellia convergence, requiring a syn-rotation, westward dipping subduction zone between these plates.

Scenario S also requires roll-back of the subduction zone below the Wrangellia superterrane in the Late Triassic-Earliest Jurassic, but southward, over  $\sim 15^{\circ}$  latitude. This roll-back must have been followed by a northward displacement over as much as 30° until the Middle Jurassic, by which time Caribbean lithosphere was forming to the south of the Wrangellia superterrane around the equator (Boschman et al., 2019). This northward Wrangellia motion occurred during activity of the Bonanza arc, and thus took place during subduction. Hence, it must have been associated with slab advance and dragging over large distances. In Scenario S, it is possible that the Socorro slab represents the lithosphere that detached below the Bonanza arc upon its cessation in the Middle Jurassic. Following this cessation, northward motion of the Wrangellia superterrane continued while the plate carrying it rotated clockwise. This rotation is faster than that reconstructed for the Farallon Plate from Pacific anomalies (e.g., Seton et al., 2012), so it requires that the Wrangellia superterrane was part of a separate plate. However, no subduction zones are required between the Wrangellia superterrane and the Farallon Plate: its clockwise rotation could have been accommodated by subduction between the Wrangellia superterrane and NAM. The systematic northward motions are consistent with the paleomagnetic evidence from accreted OPS (Courtillot et al., 1985; Tarduno et al., 1985, 1986) and North American upper plate ophiolites (Arkula et al., 2023). Subduction reactivation west of the Wrangellia superterrane is only required upon its accretion to NAM (and may thus serve as a constraint on collision age).

# 7. Conclusions

The Wrangellia superterrane is widely thought to have been part of plate(s) of the eastern Panthalassa Ocean that moved independently and converged (obliquely) with NAM during most of the Mesozoic. In this study, we reevaluate Wrangellia's paleomagnetic record and conclude with the following key points:

- 1. Our new paleomagnetic data from Upper Triassic limestones and Lower Jurassic volcanic flows of northern Vancouver Island indicate that the Bonanza arc was located either  $\sim 25-30^{\circ}$  North and rotated  $\sim 80^{\circ}$  counterclockwise (relative to present-day spin axis), or that it was located  $\sim 25-30^{\circ}$  South and rotated  $\sim 100^{\circ}$  clockwise.
- 2. New and previous Triassic to Early Jurassic paleomagnetic data of the Wrangellia superterrane are coherent in terms of inclination and declination throughout a 2,000 km-long stretch of the superterrane that lies south of the Alaskan orocline. This consistency suggests that local vertical-axis rotations linked to post-accretionary tectonics have not obscured the pre-accretionary motions of the Wrangellia superterrane.
- 3. To consider the hemispheric ambiguity of the paleomagnetic data, we present two scenarios in which the Wrangellia superterrane was located either in the southern hemisphere (Scenario S) or in the northern hemisphere (Scenario N) at ~190 Ma. From ~190 to 80 Ma, the Wrangellia superterrane either moved northward while rotating 110° clockwise relative to NAM at a north-dipping subduction zone (Scenario S) or remained at northern middle latitudes while rotating 70° counterclockwise at a south-dipping subduction zone (Scenario N). In both scenarios N and S, the accretion of the Wrangellia superterrane to the Intermontane superterrane must have taken place >1,500 km south of their present location.
- 4. The main conclusion that can be drawn from comparing these scenarios is that scenario S represents the simplest solution to transport the Wrangellia superterrane from the equator to a position alongside NAM allowing its accretion after a significant (clockwise) rotation. Moreover, Scenario S makes it possible that the

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Middle Jurassic. The  $\sim$ 5,000 km-translation in  $\sim$ 110 Ma from the southern hemisphere toward NAM is compatible with rates and directions of Jurassic oblique subduction beneath California and Cretaceous OPS motions. In contrast, scenario N requires at least two additional steps in the overall motion of the Wrangellia superterrane toward NAM: (a) after a rapid northward shift of  $\sim$ 50° in the Late Triassic to Earliest Jurassic, the superterrane moved southward relative to NAM during the Early Jurassic; (b) the superterrane rotated away from NAM during the Middle Jurassic before moving eastward to be accreted to NAM. We prefer the southern hemisphere scenario, but stress that a detailed plate kinematic restoration based on systematic analysis of Cordilleran orogenic architecture is required to fully unveil the implications of both scenarios. In addition, new paleomagnetic investigations of preserved rock records that span the current "data gap" from the Middle Jurassic to mid-Cretaceous are highly needed for improving reconstructions of the pre-accretion history of the Wrangellia superterrane.

Socorro slab (17° N) represents the lithosphere that detached below the Bonanza arc upon its cessation in the

## **Data Availability Statement**

New paleomagnetic data can be imported into the Paleomagnetism.org 2.4.0 portal from data files (.col) available in Data Sets S1 and S2, and can also be found in the Paleomagnetism.org database (Andjić et al., 2024, https://paleomagnetism.org/publication?0450afe68406c017af37b4dc8087c7bb601f2e03f8c14b9fcd64a8c4846f5963).

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